Heavy $\bar{q}q$ Free Energy and Thermodynamics in AdS/QCD

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EMMI Workshop on String Theory and Extreme Matter 15-19 March 2010, Heidelberg, Germany. H.J.Pirner & B.Galow PLB679(2009); H.J.Pirner & J.Nian NPA833(2010); B. Galow et al. 0911.0627(2090); K.Veschgini et al. 0911.1680(2009); E.Megias et al. in preparation (2010).



Motivation

Heavy $\bar{q}q$ potential at zero temperature

- Scale Invariance and Confinement
- Soft-wall model of AdS/QCD
- The 5D Einstein-dilaton model
- 3 Thermodynamics of AdS/QCD
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 - Thermodynamics
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 Free Energy
 - Polyakov Loop
 - Heavy $\bar{q}q$ free energy at $T > T_c$
 - Spatial Wilson Loops
 - Heavy $ar{m{q}}m{q}$ free energy at $T\leq T_c$

Conclusions

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Pressure of Gluodynamics

Weak Coupling Expansion and Resummed Perturbation Theory E. Braaten and A. Nieto (1996), J.O. Andersen et al (1999).



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Interaction Measure in Gluodynamics Weak Coupling Expansion and Resummed Perturbation Theory E. Braaten and A. Nieto (1996), J.O. Andersen et al (1999).



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Trace Anomaly $N_c = 3, N_f = 0$ G. Boyd et al., Nucl. Phys. B469, 419 (1996).



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Perturbation Theory and Hard Thermal Loops only yield a_{Δ} !!.

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Scale Invariance and Confinement

Consider a rectangular Wilson loop:

$$W(\mathcal{C}) = \exp\left(\mathit{ig} \int_{\mathcal{C}} \mathsf{A}_{\mu} \mathit{d} \mathsf{x}^{\mu}
ight)$$



It is related to the potential $V_{q\bar{q}}(R)$ acting between charges q and \bar{q} :

$$W(\mathcal{C}) \to \exp\left(-T \cdot V_{q\bar{q}}(R)\right)$$

Scale transformations: $T \rightarrow \lambda T$, $R \rightarrow \lambda R$, The only scale invariant solution is the Coulomb Potential:

$$V_{q\bar{q}}\sim rac{1}{F}$$

Running coupling and string tension break scale invariance:

$$V_{qar{q}}(r) = -rac{4}{3}rac{lpha_{s}(R)}{R} + \sigma R$$

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Soft-wall model of AdS/QCD

$$ds_{\text{QCD}}^2 = h(z) \cdot ds^2 = h(z) \frac{L^2}{z^2} \langle -dt^2 + d\vec{x}^2 + dz^2 \rangle.$$

• $h(z) = 1 \Longrightarrow \text{Conformal}$

• $h(z) \neq 1 \Longrightarrow$ Non conformal

Breaking of scaling invariance in QCD is given by the running coupling:

$$\Delta \equiv rac{\epsilon - 3 p}{T^4} = rac{eta(lpha_{s})}{4 lpha_{s}^2} \langle F_{\mu
u}^2
angle \,.$$

where $\beta(\alpha_s) = \mu \frac{d\alpha_s}{d\mu}$ and $\alpha_s(E) \sim 1/\log(E/\Lambda)$.

⇒Assume an ansatz for conformal invariance breaking similar to 1-loop running coupling (H.J.Pirner & B. Galow '09):

$$h(z) = rac{\log(\epsilon)}{\log(\epsilon + (\Lambda z)^2)}, \qquad z \sim rac{1}{E}.$$

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Soft-wall model of AdS/QCD

H.J.Pirner & B.Galow '09

 $W(\mathcal{C}) pprox \exp\left(-S_{
m NG}
ight) \propto \exp\left(-T \cdot V_{qar{q}}(R)
ight)$



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The 5D Einstein-dilaton model

5D Einstein-dilaton model (Gürsoy et al. '08):

$$S_{5D} = \frac{1}{16\pi G_5} \int d^5 x \sqrt{-g} \left(R - \frac{4}{3} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right) - \frac{1}{8\pi G_5} \int_{\partial M} d^4 x \sqrt{-h} \, K \, .$$

One to one relation between β -function and dilaton potential $V(\phi)$:

$$V(\phi) = V_0 \left(1 - \left(\frac{\beta(\alpha)}{3\alpha} \right)^2 \right) \exp \left[-\frac{8}{9} \int_0^\alpha \frac{\beta(a)}{a^2} da \right], \qquad \alpha = e^{\phi}.$$

Ansatz:

$$\beta(\alpha) = -b_2\alpha + \left[b_2\alpha + \left(\frac{b_2}{\bar{\alpha}} - b_0\right)\alpha^2 + \left(\frac{b_2}{2\bar{\alpha}^2} - \frac{b_0}{\bar{\alpha}} - b_1\right)\alpha^3\right]e^{-\alpha/\bar{\alpha}}.$$

• $\alpha << \bar{\alpha} \Longrightarrow$ Ultraviolet: $\beta(\alpha) \approx -b_0 \alpha^2 - b_1 \alpha^3$

• $\alpha >> \bar{\alpha} \Longrightarrow$ Infrared: $\beta(\alpha) \approx -b_2 \alpha$

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The 5D Einstein-dilaton model

$$ds^{2} = e^{\frac{4}{3}\phi}e^{2A} (dt^{2} + d\vec{x}^{2}) + e^{\frac{4}{3}\phi}\frac{12}{V_{0}}e^{2D}d\alpha^{2}$$

$$\rho(\alpha_{0}) = \int_{-\frac{\rho}{2}}^{\frac{\rho}{2}} d\sigma = \frac{12}{\sqrt{3V_{0}}}e^{-A_{0}} \cdot \int_{0}^{\alpha_{0}} \frac{e^{D-3\tilde{A}} \cdot \tilde{\alpha}^{-\frac{4}{3}}}{\sqrt{1 - \tilde{\alpha}^{-\frac{8}{3}}e^{-4\tilde{A}}}}d\alpha,$$

$$V_{Q\bar{Q}}^{\text{reg.}}(\alpha_{0}) = \frac{1}{T}S_{NG}^{\text{reg.}} = V - V_{S}$$

$$= \frac{12\alpha_{0}^{\frac{4}{3}}e^{A_{0}}}{\pi I_{S}^{2}\sqrt{3V_{0}}} \left[\int_{0}^{\alpha_{0}} d\alpha \frac{\tilde{\alpha}^{\frac{4}{3}}e^{D+\tilde{A}}}{\sqrt{1 - \tilde{\alpha}^{-\frac{8}{3}}e^{-4\tilde{A}}}} - \int_{0}^{\infty} d\alpha \tilde{\alpha}^{\frac{4}{3}} \cdot e^{D}\right]$$

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Heavy QQ potential at Short Distances

Analytical result at short distances:

$$V_{Q\bar{Q}}(\rho) = -\frac{24}{\pi V_0 \bar{J}_s^2} \frac{\alpha_0^{4/3}(\rho)}{\rho} \Big[\underbrace{0.359}_{_{\rm LO}} + \underbrace{0.533b_0\alpha_0}_{_{\rm NLO}} + \underbrace{(1.347b_0^2 + 0.692b_1)\alpha_0^2}_{_{\rm NNLO}}\Big]$$

where the running coupling is

$$\alpha_{0}(\rho) = \left[b_{0}\log\left(\frac{4.57}{\rho\sqrt{V_{0}}}\right) + \frac{b_{1}}{b_{0}}\log\left(b_{0}\log\left(\frac{4.57}{\rho\sqrt{V_{0}}}\right)\right)\right]^{-1} + \mathcal{O}\left(\log^{-3}\right)$$
$$|\rho V_{00}(\rho)|$$

0.40 NNLO 0.35 0.30 NLO 0.25 LO 0.20 0.08 0.10 0.12 0.14 0.16 0.18 0.20 ρ [GeV⁻¹] 0.15

Comparison with $PT \Longrightarrow$

$$V_0 \overline{I}_s^2 = 1.31$$

(0)

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Heavy $\bar{Q}Q$ potential and running coupling



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Dilaton potential and warp factor



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Confinement and 'good' IR singularity

Effective Schrödinger potentials for glueballs 0⁺⁺ and 2⁺⁺:



• Confining theory \implies 1.5 < b_2

• IR singularity repulsive to physical modes $\implies b_2 < 2.37$

Best choice of parameters:

 $b_2 = 2.3$, $\bar{\alpha} = 0.45$, $V_0 = -0.623 \,\mathrm{GeV}^2$, $\bar{l}_s = 1.45 \,\mathrm{GeV}^{-1}$

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Schwarzschild black hole

General Relativity with no source \implies Einstein-Hilbert action

$$S_{EH}=rac{1}{16\pi G_D}\int d^Dx\sqrt{-g}\ R\,,\qquad R=g^{\mu
u}R_{\mu
u}$$

Classical solution $rac{\delta}{\delta g_{\mu
u}} \Longrightarrow$ Einstein equations

$$E_{\mu
u}\equiv R_{\mu
u}-rac{1}{2}g_{\mu
u}R=0 \Longrightarrow_{ ext{spherical}}R_{\mu
u}=0$$

Schwarzschild solution in spherical coordinates (1915):

$$ds^{2} = g_{\mu\nu}dx^{\mu}dx^{\nu} = -f(r)dt^{2} + \frac{1}{f(r)}dr^{2} + r^{2}d\Omega_{2}^{2}, \qquad f(r) = 1 - \frac{r_{h}}{r}$$

 r_h is the horizon. Not physical singularity: $R^{\mu\nu\rho\sigma}R_{\mu\nu\rho\sigma} = 12\frac{r_h^{\epsilon}}{r^6}$. Large distance limit: $g_{tt}(r) \underset{r \to \infty}{\sim} -(1 + 2V_{Newton}(r)) \Longrightarrow r_h = 2G_4M$.

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Black hole thermodynamics

$$Z = \operatorname{Tr}\left(\mathbf{e}^{-eta H}
ight), \qquad eta = rac{1}{T}$$

Periodicity in euclidean time ($\tau = it$): $\Phi(\tau + \beta) = \Phi(\tau)$

• Regularity: Expansion around the horizon $r = r_h(1 + \rho^2)$:

$$ds^{2} = 4r_{h}^{2}\left(d\rho^{2} + \rho^{2}\underbrace{\left(\frac{d\tau}{2r_{h}}\right)^{2}}_{d\theta^{2}} + \frac{1}{4}d\Omega_{2}^{2}\right)$$

$$\implies \text{Periodicity: } \frac{\tau}{2r_h} \rightarrow \frac{\tau}{2r_h} + 2\pi \implies \tau \rightarrow \tau + 4\pi r_h =: \tau + \beta$$
$$T = \frac{1}{8\pi MG_4}$$

Thermodynamics interpretation of black holes:

$$dM = TdS \Longrightarrow S = \int \frac{dM}{T} = 4\pi G_4 M^2$$

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Black hole thermodynamics

Area of the event horizon: $A = 4\pi r_h^2 = 16\pi (G_4 M)^2$. Bekenstein-Hawking entropy formula:

$$S = rac{\mathcal{A}}{4G_D}$$

Reissner-Nordström black hole:

$$S_{\text{Einstein}-Maxwell} = \int d^{D}x \sqrt{-g} \left(\frac{1}{16\pi G_{D}}R - \frac{1}{4}F_{\mu\nu}^{2}\right)$$
Solution: $f(r) = 1 - \frac{2G_{4}M}{r} + \frac{G_{4}Q^{2}}{r^{2}}$.
$$Q^{2} > G_{4}M^{2} \text{ no singularity}$$

$$Q^{2} = G_{4}M^{2} \text{ extremal, T=0}$$

$$Q^{2} < G_{4}M^{2} \text{ two singularities}$$

$$Q^{2} = 0 \text{ one singularity} = A = 0$$

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The 5D Einstein-dilaton model

$$S_{5D} = \frac{1}{16\pi G_5} \int d^5 x \sqrt{-g} \left(R - \frac{4}{3} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right) - \frac{1}{8\pi G_5} \int_{\partial M} d^4 x \sqrt{-h} K$$

Finite temperature solutions (E. Kiritsis et al. JHEP (2009) 033):

• Thermal gas solution (confined phase):

$$ds_{\rm th}^2 = b_0^2(z)(-dt^2 + d\vec{x}^2 + dz^2)\,, \qquad t \sim t + i\beta$$

• Black hole solution (deconfined phase):

$$ds_{\rm BH}^2 = b^2(z) \left[-\frac{f(z)}{dt^2} + d\vec{x}^2 + \frac{dz^2}{f(z)} \right]$$

In the UV ($z \simeq 0$): flat metric $b(z) \simeq L/z$ and f(0) = 1. There exists an horizon $f(z_h) = 0$. Regularity at the horizon $\implies T = \frac{|\dot{f}(z_h)|}{4\pi}$.

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The 5D Einstein-dilaton model

Einstein equations $\frac{\delta}{\delta g_{\mu\nu}}$:

$$\underbrace{\begin{pmatrix} R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R \\ E_{\mu\nu} \\ (a) & \frac{\ddot{f}}{\dot{f}} + 3\frac{\dot{b}}{b} = 0, \Longrightarrow f(z) = 1 - \frac{\int_{0}^{z} \frac{d\bar{z}}{b(\bar{z})^{3}}}{\int_{0}^{z_{h}} \frac{d\bar{z}}{b(\bar{z})^{3}}} \\ (b) & 6\frac{\dot{b}^{2}}{b^{2}} - 3\frac{\ddot{b}}{b} = \frac{4}{3}\dot{\phi}^{2}, \\ (c) & 6\frac{\dot{b}^{2}}{b^{2}} + 3\frac{\ddot{b}}{b} + 3\frac{\dot{b}}{f} = \frac{b^{2}}{f}V(\phi) \end{aligned}$$

Conformal solution:

$$V(\phi) = \frac{12}{L^2}, \quad \dot{\phi} = 0 \Longrightarrow b(z) = \frac{L}{z}, \quad f(z) = 1 - \left(\frac{z}{z_h}\right)^4, \quad T = \frac{1}{\pi z_h}$$

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$$b^2(z) = e^{-\frac{4}{3}\phi(z)}\frac{L^2}{z^2}h(z)$$
, Input: $h(z) = \frac{\log(\epsilon)}{\log(\epsilon + (\Lambda z)^2)}$ Pirner&Galow'09.



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Thermodynamics

Postulate: Entropy of gauge theories is equal to the Bekenstein-Hawking entropy of their string duals.

$$S(T) = rac{\mathcal{A}(z_h)}{4G_5} = rac{V_3 b^3(z_h)}{4G_5} = V_3 s_0 rac{h^{\frac{3}{2}}(z_h)}{z_h^3}, \qquad z_h = rac{1}{\pi T}$$

High temperature limit: $s(T) \underset{T \to \infty}{\sim} s_0 \pi^3 T^3 = \frac{32}{45} \pi^2 T^3 =: s_{ideal}(T)$ One can compute all the thermodynamics quantities:

$$\mathbf{s}(T) = rac{d}{dT} \mathbf{p}(T), \qquad \Delta(T) \equiv rac{\epsilon - 3\mathbf{p}}{T^4} = rac{\mathbf{s}}{T^3} - rac{4\mathbf{p}}{T^4}.$$

One can choose several warp factors:

- Andreev & Zakharov '07: $h_A(z) = e^{1/2cz^2}$.
- Pirner & Galow '09: $h_P(z) = \frac{\log(\epsilon)}{\log(\epsilon + (\Lambda z)^2)}, \quad \epsilon = \frac{l_s^2}{L^2}.$

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Including dilaton dynamics:



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Polyakov Loop

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Polyakov Loop

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Polyakov Loop

$$L(T) := \langle P \rangle = \int DX e^{-S_w} \underset{\text{semiclassically}}{\Longrightarrow} \langle P \rangle = \sum_i w_i e^{-S_i} \simeq w_0 e^{-S_0}$$

Nambu-Goto Action:

$$\begin{split} S_{\rm NG} &= \frac{1}{2\pi l_{\rm S}^2} \int d\sigma d\tau \sqrt{\det h_{ab}} = \frac{1}{2\pi l_{\rm S}^2} \int d\sigma d\tau \sqrt{\det g_{\mu\nu} \partial_a X^\mu \partial_b X^\nu} \,, \\ \mu, \nu &= t, \vec{x}, z \qquad a, b = \sigma, \tau \end{split}$$

Modified AdS₅-metric at finite temperature:

$$ds_E^2 = rac{L^2}{z^2}h(z)\left(f(z)dt^2 + d\vec{x}^2 + rac{dz^2}{f(z)}
ight), \qquad f(z) = 1 - \left(rac{z}{z_h}
ight)^4$$

Static configurations: $\tau = t$, $\sigma = z$

$$h_{ab} = \frac{L^2}{z^2} h(z) \begin{pmatrix} f & 0\\ 0 & \frac{1}{f} + \dot{x}^2 \end{pmatrix}.$$

Polyakov Loop

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Polyakov Loop

The NG action writes:

$$S_{\rm NG} = rac{L^2}{2\pi l_s^2} \int_0^{1/T} d au \int_0^{z_h} dz rac{h(z)}{z^2} \sqrt{1 + \dot{x}^2 f(z)} \, ,$$

Equation of motion for *x*:

$$\frac{\partial}{\partial z} \left[\frac{h(z)}{z^2} \frac{\dot{x}f}{\sqrt{1 + \dot{x}^2 f}} \right] = 0, \qquad \text{Boundary cond.: } x(0) = x(z_h) \equiv x_0$$

Solution: $x = x_0 =$ **constant**.

$$S_{\mathrm{NG}}^{\mathrm{reg}}=-rac{1}{2\epsilon}+rac{1}{2\pi\epsilon T}\int_{0}^{z_{h}}dzrac{h(z)-1}{z^{2}}\,.$$

Polyakov Loop

Heavy $\bar{q}q$ free energy at $T > T_{C}$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T < T_{C}$

Polyakov Loop



Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_C$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \le T_C$

Why are warp factors so good?

 $h_A(z)$ and $h_P(z)$ both describe very well lattice results. This is because they include power corrections:

$$h_P(z) = \frac{\log(\epsilon)}{\log(\epsilon + (\Lambda z)^2)} = 1 + \frac{\Lambda^2}{\epsilon |\log \epsilon|} z^2 + \frac{(2 + \log \epsilon)\Lambda^4}{2\epsilon^2 \log^2 \epsilon} z^4 + \mathcal{O}(z^6).$$

They are related to condensates, starting from dimension 2.

- Dim 2 condensate: $C_2 \equiv g^2 \langle A_{0,a}^2 \rangle = \frac{2N_c}{\pi^2} \frac{\Lambda^2}{\epsilon^{2} |\log \epsilon|} = (0.50 \,\text{GeV})^2$,
- Polyakov loop:

$$L \simeq e^{-S_0} = \exp\left(c_0 - \frac{C_2}{4N_cT^2} - \frac{C_4}{16N_c^2T^4} + \cdots\right),$$

• Trace anomaly:

$$\frac{\epsilon - 3p}{T^4} = \frac{33}{4\pi} \alpha_s \frac{\mathcal{C}_2}{T^2} + \mathcal{O}\left(\frac{\mathcal{C}_4}{T^4}\right)$$

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Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

Perturbation theory

But they don't fulfill UV behaviour given by perturbation theory.

$$L = 1 + \underbrace{\frac{4}{3}\sqrt{\pi}\alpha_{s}^{3/2}}_{\sim 1/\log^{3/2}(T/\Lambda)} + \underbrace{(2\log\alpha_{s} + 3 + 2\log\pi)\alpha_{s}^{2}}_{\sim 1/\log^{2}(T/\Lambda)} + \mathcal{O}(\alpha_{s}^{5/2}) \quad \text{Gava '81}$$
$$\frac{\epsilon - 3\rho}{T^{4}} = \underbrace{\frac{11}{3}\alpha_{s}^{2}}_{\sim 1/\log^{2}(T/\Lambda)} + \mathcal{O}(\alpha_{s}^{5/2})$$

One can try to construct the corresponding h(z):

$$h(z) \sim \log^{-n}(\Lambda z), \qquad n > 0$$

For the Polyakov loop:

$$h(z) = 1 + \epsilon \frac{16\sqrt{2}}{33\sqrt{11}} \pi^2 \left(\frac{1}{\log^{3/2}(1/(\Lambda z))} - \frac{3}{2} \frac{1}{\log^{5/2}(1/(\Lambda z))} \right) \xrightarrow[z \to 0]{} 1$$

 $\begin{array}{l} \textbf{Polyakov Loop} \\ \textbf{Heavy } \bar{q}q \text{ free energy at } T > T_{\mathcal{C}} \\ \textbf{Spatial Wilson Loops} \\ \textbf{Heavy } \bar{q}q \text{ free energy at } T \leq T_{\mathcal{C}} \end{array}$

A different approximation

Different approximation: choose the form of the dilaton potential (Kiritsis '09, Kajantie '09, Megias '09).

$$\beta(\alpha) = -\beta_0 \alpha^2 - \beta_1 \alpha^3 + \dots \Longrightarrow$$
$$\implies V(\alpha = e^{\phi}) = \frac{12}{L^2} \left(1 + \frac{8}{9} \beta_0 \alpha + \left(\frac{23}{81} + \frac{4\beta_1^2}{9\beta_0^2} \right) (\beta_0 \alpha)^2 + \dots \right)$$

Kajantie et al. arXiv:0905.2032 (2009), spatial string tension:

$$V(r) = \sigma_s r$$
, $\sigma_s(T) = \frac{1}{2\pi \alpha'} b^2(z_h) \alpha^{4/3}(z_h)$

This is in contradiction with QCD: $\sigma_{\text{QCD}} \sim T^2 \alpha_s^2(T)$. • Megías et al. (2010) compute the Polyakov loop:

$$L(T) = \exp\left(c_0 + \frac{C_*}{2\epsilon}\alpha^{\frac{4}{3}}(\mathbf{z}_h) + \mathcal{O}(\alpha^{\frac{7}{3}})\right)$$

In contradiction with PT: $L_{PT}(T) = \exp\left(\frac{4}{3}\sqrt{\pi\alpha^{\frac{3}{2}}} + \mathcal{O}(\alpha^{2})\right)$.

Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

Issues

Motivation

- Heavy $\bar{q}q$ potential at zero temperature
 - Scale Invariance and Confinement
 - Soft-wall model of AdS/QCD
 - The 5D Einstein-dilaton model
- 3 Thermodynamics of AdS/QCD
 - Black Holes
 - The 5D Einstein-dilaton model at finite temperature
 - Thermodynamics

Heavy q q q Free Energy

- Polyakov Loop
- Heavy $\bar{q}q$ free energy at $T > T_c$
- Spatial Wilson Loops
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Conclusions

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Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

Heavy $\bar{q}q$ free energy at $T > T_c$

$$\mathbf{e}^{-eta F_{qar{q}}} = \left\langle \Omega\left(\mathbf{x} = rac{d}{2}
ight) \Omega^{\dagger}\left(\mathbf{x} = -rac{d}{2}
ight)
ight
angle \simeq \mathbf{e}^{-S_{
m NG}}$$

The Nambu-Goto action writes:

$$S_{\rm NG} = \frac{1}{2\pi\epsilon} \int_0^\beta d\tau \int_{-d/2}^{d/2} dx \frac{h(z)}{z^2} \sqrt{f(z) + \dot{z}^2} \,, \quad X^\mu(\tau, x) = (\tau, x, 0, 0, z(x)) \,.$$

and the heavy $\bar{q}q$ free energy:

$$\begin{aligned} F_{\bar{q}q} &= T \cdot S_{NG}^{reg} = \frac{1}{\pi \epsilon} \left[\int_{0}^{z_0} \frac{dz}{z^2} \left(\frac{h^2(z)f(z)}{\sqrt{h^2(z)f^2(z) - f(z)\frac{z^4}{C^4}}} \right) - \int_{0}^{z_h} dz \frac{h(z)}{z^2} \right] \\ d &= 2 \int_{0}^{z_0} dz \frac{z^2}{\sqrt{C^4 h^2(z)f^2(z) - f(z)z^4}}, \qquad z_0 = z(x=0) \end{aligned}$$

(1)

Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_C$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_C$

Heavy $\bar{q}q$ free energy at $T > T_c$



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Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

Heavy $\bar{q}q$ free energy at $T > T_c$

Debye mass:

$$F_{ar{q}q}(d) \sim rac{e^{-m_D\dot{d}}}{d} \Longrightarrow m_D \sim rac{1}{d_{
m screening}}$$



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Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \le T_c$

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Spatial Wilson Loops

Rectangular Wilson loop in (x,y) plane:

$$\langle W(\mathcal{C}) \rangle = \left\langle \exp\left(ig \int_{\mathcal{C}} A_{\mu} dx^{\mu}\right) \right\rangle^{l_{y} \to \infty} e^{-l_{y} \cdot V(d)},$$

$$V(d) \stackrel{d \to \infty}{\simeq} \sigma_{s} \cdot d$$

$$S_{\mathrm{NG}} = \frac{L^{2}}{2\pi l_{s}^{2}} l_{y} \int_{-l_{y}/2}^{l_{x}/2} dx \frac{h(z)}{z^{2}} \sqrt{1 + \frac{(\partial_{x} z)^{2}}{f(z)}}, \quad X(x, y) = (x, y, 0, 0, z(x))$$

Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_C$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_C$

Spatial Wilson Loops



Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

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Heavy qq Free Energy

- Polyakov Loop
- Heavy $\bar{q}q$ free energy at $T > T_c$
- Spatial Wilson Loops
- Heavy $\bar{q}q$ free energy at $T \leq T_c$

Conclusions

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Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \le T_c$

Heavy $\bar{q}q$ free energy at $T \leq T_c$

K.Veschgini, E.Megías, J.Nian & H.J.Pirner, arXiv:0911.1680 ('09).

$$\mathbf{e}^{-eta F_{qar{q}}(ec{d},T)} = rac{1}{N_c^2} \Big\langle \mathrm{tr}_c \Omega(rac{ec{d}}{2}) \, \mathrm{tr}_c \Omega^\dagger(-rac{ec{d}}{2}) \Big
angle pprox \mathbf{e}^{-S_{\mathrm{NG}}}$$



Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_c$

Heavy $\bar{q}q$ free energy at $T \leq T_c$

$$S_{\rm NG} = \frac{1}{2\pi l_s^2} \int_0^{2\pi} d\phi \int_{-d/2}^{d/2} dx \frac{L^2 h(z)}{z^2} r \sqrt{1 + (z')^2 + (r')^2}$$

Euler-Lagrange equations:

(a)
$$k = \frac{h(z) \cdot r}{z^2} \frac{1}{\sqrt{1 + (z')^2 + (r')^2}}$$

(b) $r'' - \frac{h^2(z) \cdot r}{k^2 z^4} = 0$
(c) $z'' - \frac{h(z) \cdot r^2 \cdot (z \partial_z h(z) - 2h(z))}{k^2 z^5} = 0$

Boundary conditions:

$$r(\pm d/2) = R = \frac{\beta}{2\pi} = \frac{1}{2\pi T}, \qquad z(\pm d/2) = 0.$$

Eugenio Megías Heavy qq free energy a

Heavy $\bar{q}q$ free energy and Thermodynamics in AdS/QCD

Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_c$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \le T_c$

Heavy $\bar{q}q$ free energy at $T \leq T_c$

Numerical solutions only in a limited range. No minimal surface \implies classical approximation not valid.



Polyakov Loop Heavy $\bar{q}q$ free energy at $T > T_C$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T \leq T_C$

Heavy $\bar{q}q$ free energy at $T \leq T_c$



Eugenio Megías Heavy qq free energy and Thermodynamics in AdS/QCD

Polyakov Loop Heavy $\bar{a}a$ free energy at $T > T_C$ Spatial Wilson Loops Heavy $\bar{q}q$ free energy at $T < T_{C}$

Heavy $\bar{q}q$ free energy at $T < T_c$



Other thermodynamics quantities:

Entropy:

$$S_{\bar{q}q} = -rac{\partial F_{\bar{q}q}}{\partial T} = rac{7.46}{\mathrm{fm}}d,$$

Inner energy:

$$\mathsf{E}_{\bar{q}q} = \mathsf{F}_{\bar{q}q} + \mathsf{T} \cdot \mathsf{S}_{\bar{q}q} = \frac{-0.48}{d} + \frac{5.84}{\mathrm{fm}^2} \mathsf{d},$$

Conclusions:

- The non-perturbative behaviour of QCD near and above *T_c* is characterized by power corrections in *T*. These power corrections are high energy trace of non-perturbative low energy effects.
- AdS-QCD serves as a powerful tool to study the non perturbative regime of QCD at zero temperature (large distances) and finite temperature (close to the phase transition). We consider conformal breaking warp factors which naturally describe these power corrections.
- We describe at the same time the equation of state of QCD, and the heavy q
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 free energy.

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